

Conservative High-Order Finite-Difference Schemes for Low-Mach Number Flows

F. Nicoud¹

Center for Turbulence Research, Stanford University, Stanford, California 94305-3030

E-mail: nicoud@cerfacs.fr

Received January 18, 1999; revised November 8, 1999

Three finite-difference algorithms are proposed to solve a low-Mach number approximation for the Navier–Stokes equations. These algorithms exhibit fourth-order spatial and second-order temporal accuracy. They are dissipation-free, and thus well suited for DNS and LES of turbulent flows. The key ingredient common to each of the methods presented is a Poisson equation with variable coefficient that is solved for the hydrodynamic pressure. This feature ensures that the velocity field is constrained correctly. It is shown that this approach is needed to avoid violation of the conservation of kinetic energy in the inviscid limit which would otherwise arise through the pressure term in the momentum equation. An existing set of finite-difference formulae for incompressible flow is generalized to handle arbitrary large density fluctuations with no violation of conservation through the non-linear convective terms. An algorithm which conserves mass, momentum, and kinetic energy fully is obtained when an approximate equation of state is used instead of the exact one. Results from a model problem are used to show both spatial and temporal convergence rates and several test cases are presented to illustrate the performance of the algorithms. © 2000 Academic Press

Key Words: low-Mach number; finite difference; staggered mesh; density gradient; conservative scheme; Poisson equation.

1. INTRODUCTION

Large Eddy simulation (LES) and/or direct numerical simulation (DNS) can provide detailed information about turbulent flows that may be difficult to obtain experimentally. However, for the particular class of flow with low-Mach number M and strong density variation, the classical compressible Navier–Stokes equations are not well suited for computation. The small time step limitation dictated by numerical stability requirements of explicit methods would require excessive computing times to solve practical flow problems.

¹ Permanent address: CERFACS, 42, Av. Gaspard Coriolis, 31057 Toulouse Cedex 1, France.

Indeed the sound waves move much faster than entropy or vorticity waves when $M \ll 1$. At the same time, in flows where the dominating mechanism is free, forced or mixed convection, the acoustical mode of energy carries only a small fraction of the energy present in the fluctuating part of the flow. These observations led several authors [1–4] to propose a set of low-Mach number equations which do not contain acoustic waves but can still describe the entropy and vorticity modes as well as compressibility due to exothermicity of chemical reactions. A fractional-step method is used most often, the pressure field being obtained by solving a Poisson equation with the time derivative of the density field as part of the source term [3, 5] destabilizing part of the algorithm [6]. Even-ordered finite difference approximations to this derivative were found to be more stable but density ratios larger than 3 are difficult to compute. Sandoval (reported in [6]) found that by decreasing the Reynolds number, larger variations in density could be achieved. Larger density ratios seem computable by using a predictor-corrector time-stepping algorithm in which the predictor uses a second-order Adams–Bashforth time integration scheme and the corrector relies on a quasi-Crank–Nicolson integration with the inversion of a pressure Poisson equation at each step [7, 8].

As far as incompressible Navier–Stokes equations are concerned, experience has shown that the kinetic energy must be conserved if a stable and dissipation-free numerical method is sought. Indeed, such a property ensures that the sum of the square of the velocities cannot grow, even through non-linear interactions between modes. As a consequence, a numerical scheme which conserves kinetic energy cannot be unstable. Moreover, it makes unnecessary the use of numerical stabilization through up-winding which is known to introduce too much artificial damping in DNS/LES computations. Morinishi *et al.* [9] developed a set of fully conservative (mass, momentum, and kinetic energy) high order schemes for incompressible flow. However, none of the numerical studies on low-Mach number flows cited above addressed this issue. In the present study it is shown that the global conservation of kinetic energy is a common feature of incompressible and low-Mach number flows in the inviscid limit. A nearly conservative fourth-order finite difference scheme is proposed in which one solves a Poisson equation with variable coefficient for pressure. Also, this algorithm makes use of a generalization of the Morinishi’s finite difference formulae for variable density flow. These two ingredients lead to algorithms which are well suited for LES and/or DNS computations. In particular:

- no numerical dissipation from the spatial discretization is used to stabilize the computation; and,
- they can handle density ratios much larger than 3.

The low-Mach number approximation and the numerical method are discussed in Sections 2 and 3. An error analysis is conducted in Section 4 and some numerical experiments are discussed in Section 5 to show the potential of the algorithm.

2. LOW-MACH NUMBER APPROXIMATION

In compressible flows, a natural parameter to measure the effects of compressibility is the ratio of the dynamic to the thermodynamic pressure, viz. γM^2 . To derive the low-Mach number equations, one expands the dependent variables as a power series in $\epsilon = \gamma M^2$, which is a small parameter (see [1, 3, 6] for details on the derivation of the low-Mach number equations). Substituting these expansions into the compressible Navier–Stokes equations

and collecting the lowest order terms in ϵ yields

$$\frac{\partial \rho}{\partial t} + \frac{\partial \rho u_j}{\partial x_j} = 0 \quad (1)$$

$$\rho \frac{\partial u_i}{\partial t} + \rho u_j \frac{\partial u_i}{\partial x_j} = -\frac{\partial P}{\partial x_i} + \frac{1}{R_e} \frac{\partial \tau_{ij}}{\partial x_j} \quad (2)$$

$$\rho C_p \frac{\partial T}{\partial t} + \rho C_p u_j \frac{\partial T}{\partial x_j} = \frac{1}{R_e P_r} \frac{\partial q_j}{\partial x_j} + \frac{\gamma - 1}{\gamma} \frac{dP_o}{dt}. \quad (3)$$

As body forces have no impact on the numerical methods presented in Section 3, they have not been included in Eq. (2). All the variables are normalized using the reference state $\rho^{\text{ref}}, u^{\text{ref}}, T^{\text{ref}} = P_o^{\text{ref}}/\rho^{\text{ref}}, C_p^{\text{ref}} = C_p^*(T^{\text{ref}}), \mu^{\text{ref}} = \mu^*(T^{\text{ref}})$, and $k^{\text{ref}} = k^*(T^{\text{ref}})$ where the superscript * represents dimensional quantities. Also $R_e = \rho^{\text{ref}} u^{\text{ref}} L^{\text{ref}}/\mu^{\text{ref}}$ and $P_r = \mu^{\text{ref}} C_p^{\text{ref}}/k^{\text{ref}}$ are the Reynolds and the Prandtl number while γ is the ratio of specific heats at the reference state. u_i, ρ, T, k, μ , and C_p stand for the non-dimensionalized velocity vector, density, temperature, thermal conductivity, dynamic viscosity, and specific heat. $\tau_{ij} = \mu(\partial u_i/\partial x_j + \partial u_j/\partial x_i - (2/3)\delta_{ij}(\partial u_k/\partial x_k))$ and $q_j = k(\partial T/\partial x_j)$ are the viscous stress tensor and the heat flux vector, respectively. Moreover, P may be interpreted as the hydrodynamic pressure. In the low-Mach number approximation, the thermodynamic pressure P_o only depends on time and must be computed if it is not constant. The equation of state is simply

$$P_o = \rho T. \quad (4)$$

Since density is uniquely determined by the temperature (and the thermodynamic pressure which is constant in space), the energy equation acts as a constraint which is enforced by the hydrodynamic pressure. Combining Eqs. (1), (3), and (4), this constraint is

$$\frac{\partial u_i}{\partial x_i} = \frac{1}{P_o(t)C_p} \left[\frac{1}{R_e P_r} \frac{\partial}{\partial x_j} \left(k \frac{\partial T}{\partial x_j} \right) + \left(\frac{\gamma - 1}{\gamma} - C_p \right) \frac{dP_o}{dt} \right]. \quad (5)$$

Integrating over the flow domain V leads to the following ODE for the thermodynamic pressure in a closed system:

$$\frac{dP_o}{dt} = \frac{1}{\int_V \left(\frac{\gamma-1}{\gamma} - C_p \right) dV} \left[\frac{1}{R_e P_r} \int_V \frac{\partial}{\partial x_j} \left(k \frac{\partial T}{\partial x_j} \right) dV + P_o(t) \int_V u_i \frac{\partial C_p}{\partial x_i} dV \right]. \quad (6)$$

Since $\int_V (\partial/\partial x_j) (k(\partial T/\partial x_j)) dV = \int_S k(\partial T/\partial x_j) dS_j$, this relation expresses how the rate of change of the mean pressure is affected by the heat flux through the surface S of the domain V and the gradients of heat capacity of the gas. In many practical applications such as piston engines and internal ballistics the fluid may be considered as a calorifically perfect gas so that $C_p = 1$ and the time derivative of P_o is simply

$$\frac{dP_o}{dt} = \frac{\gamma}{V} \frac{1}{R_e P_r} \int_V \frac{\partial}{\partial x_j} \left(k \frac{\partial T}{\partial x_j} \right) dV = \frac{1}{V} \frac{1}{R_e P_r} \int_S k \frac{\partial T}{\partial x_j} dS_j. \quad (7)$$

Thus the constraint on the velocity field becomes

$$\frac{\partial u_i}{\partial x_i} = \frac{1}{P_o(t)R_e P_r} \left[\frac{\partial}{\partial x_j} \left(k \frac{\partial T}{\partial x_j} \right) - \frac{1}{V} \int_V \frac{\partial}{\partial x_j} \left(k \frac{\partial T}{\partial x_j} \right) dV \right]. \quad (8)$$

Note that the numerical methods presented in Section 3 remain usable even if the fluid is not considered as being a calorifically perfect gas. If the system considered is open, then the thermodynamic pressure is set by atmospheric conditions. If it is closed then the amount of mass in it, M_0 , is constant over time so that by integrating the equation of state over the whole domain one obtains the following expression for the thermodynamic pressure:

$$P_o(t) = \frac{M_0}{\int_V (1/T) dV}. \quad (9)$$

Note that in the limit of an inviscid flow of a calorifically perfect gas the thermodynamic pressure remains constant over time (from Eq. (7)) and the velocity field is divergence-free (from Eq. (8)).

The solution (ρ, u_i, T, P, P_o) is completely described by Eqs. (1)–(4) and (7). The constraint Eq. (8) should also be satisfied since it is a linear combination of Eqs. (1), (3), and (4).

3. NUMERICAL METHOD

The numerical method chosen for solving the variable density momentum and temperature equations is a generalization of a fully conservative fourth order spatial scheme developed for incompressible flows on staggered grids by Morinishi *et al.* [9]. A scheme to solve the momentum equations in non-conservative form is described in the following Subsection 3.1. A scheme with “good” conservative properties is discussed in Subsection 3.2. Both of these algorithms involve a variable coefficient Poisson equation for the pressure. For completeness, Subsection 3.3 presents an alternative formulation in which the pressure is obtained, as proposed in most of the previous studies, through a Poisson equation with constant coefficient and approximate source term.

3.1. Scheme in Non-conservative Form: AdvSC

For a uniform mesh, the advective term in the momentum equation, Eq. (2), is discretized as

$$\begin{aligned} \rho u_j \frac{\partial u_i}{\partial x_j} \equiv & \frac{9}{8} \left(\frac{9 \overline{\rho^{(4j)} u_j}^{1x_i}}{8 \overline{\rho^{(4j)} u_j}^{1x_i}} - \frac{1 \overline{\rho^{(4j)} u_j}^{3x_i}}{8 \overline{\rho^{(4j)} u_j}^{3x_i}} \right) \frac{\delta_1 u_i}{\delta_1 x_j} \\ & - \frac{1}{8} \left(\frac{9 \overline{\rho^{(4j)} u_j}^{1x_i}}{8 \overline{\rho^{(4j)} u_j}^{1x_i}} - \frac{1 \overline{\rho^{(4j)} u_j}^{3x_i}}{8 \overline{\rho^{(4j)} u_j}^{3x_i}} \right) \frac{\delta_3 u_i}{\delta_3 x_j}, \end{aligned} \quad (10)$$

where the finite-difference operator with stencil n acting on ϕ with respect to x_i is defined as

$$\frac{\delta_n \phi}{\delta_n x_i} = \frac{\phi(x_i + nh_i/2) - \phi(x_i - nh_i/2)}{nh_i} \quad (11)$$

and the interpolation operator with stencil n acting on ϕ in the x_i direction is

$$\bar{\phi}^{nx_i} = \frac{\phi(x_i + nh_i/2) + \phi(x_i - nh_i/2)}{2}. \quad (12)$$

$\rho^{(4j)}$ is a fourth order interpolation of ρ in the x_j direction,

$$\rho^{(4j)} = \frac{9}{8}\bar{\rho}^{1x_j} - \frac{1}{8}\bar{\rho}^{3x_j}. \quad (13)$$

When the density is constant, Eq. (10) reduces to the advective form (Adv.-S4) in Morinishi *et al.* [9]. The pressure term in Eq. (2) is discretized by

$$\frac{\partial P}{\partial x_i} \equiv (\text{Pres.})_i = (\nabla_d P)_i = \frac{9}{8} \frac{\delta_1 P}{\delta_1 x_i} - \frac{1}{8} \frac{\delta_3 P}{\delta_3 x_i} \quad (14)$$

and the discrete divergence operator is defined consistently,

$$\frac{\partial u_i}{\partial x_i} \equiv \nabla_d \cdot (u_i) = \frac{9}{8} \frac{\delta_1 u_i}{\delta_1 x_i} - \frac{1}{8} \frac{\delta_3 u_i}{\delta_3 x_i}. \quad (15)$$

The viscous terms in Eq. (2) are written using the generic form

$$\begin{aligned} \frac{\partial}{\partial x_j} \left(\mu \frac{\partial u_i}{\partial x_j} \right) &\equiv \frac{9}{8} \frac{\delta_1}{\delta_1 x_j} \left[\mu^{(4i),(4j)} \left(\frac{9}{8} \frac{\delta_1 u_i}{\delta_1 x_j} - \frac{1}{8} \frac{\delta_3 u_i}{\delta_3 x_j} \right) \right] \\ &\quad - \frac{1}{8} \frac{\delta_3}{\delta_3 x_j} \left[\mu^{(4i),(4j)} \left(\frac{9}{8} \frac{\delta_1 u_i}{\delta_1 x_j} - \frac{1}{8} \frac{\delta_3 u_i}{\delta_3 x_j} \right) \right]. \end{aligned} \quad (16)$$

Note that μ is successively interpolated in the i and j -direction (applying Eq. (13) to μ once in i , once in j) to give the fourth-order interpolation $\mu^{(4i),(4j)}$. The advective term for the temperature is discretized as

$$\rho u_j \frac{\partial T}{\partial x_j} \equiv \frac{9}{8} \overline{\rho^{(4j)} u_j} \frac{\delta_1 T}{\delta_1 x_j} - \frac{1}{8} \overline{\rho^{(4j)} u_j} \frac{\delta_3 T}{\delta_3 x_j}. \quad (17)$$

A semi-implicit time marching algorithm is used in which the diffusion terms in the wall normal direction are treated implicitly with a Crank–Nicolson scheme, while a third order Runge–Kutta or second order Adams–Bashforth scheme is used for all other terms. The temperature equation is advanced first so that ρ^{n+1} is known via the state equation $\rho = P_o/T$, where P_o is first assessed using Eq. (9) written at time $n + 1$. Then a fractional step method is used to solve the momentum equation,

$$\begin{aligned} \rho^{(4i),n+1} \frac{u_i^{n+1} - u_i^n}{\Delta t} &= \rho^{(4i),n+1} \frac{u_i^{n+1} - \hat{u}_i}{\Delta t} + \rho^{(4i),n+1} \frac{\hat{u}_i - u_i^n}{\Delta t} \\ &= \beta_k (I^{n+1} + I^n) + \gamma_k E^n + \zeta_k E^{n-1} - 2\beta_k \nabla_d P^n - 2\beta_k \nabla_d \delta P^{n+1}, \end{aligned} \quad (18)$$

where I and E represent all the spatial implicit and explicit terms except for the pressure at n and the pressure update $\delta P^{n+1} = P^{n+1} - P^n$. The parameters β_k , γ_k , and ζ_k ($k = 1, 3$) can be chosen so that the mixed Runge–Kutta/Crank–Nicolson (RK-CN) time stepping is recovered after the third sub-step (see Spalart [10]). One can also choose their values so that the mixed Adams–Bashforth/Crank–Nicolson (AB-CN) time stepping is obtained at each sub-step. The values for the coefficients β_k , γ_k , and ζ_k are given in Table I.

TABLE I
Numerical Parameters for Time Stepping
RK-CN and AB-CN

Sub-step	β_k	γ_k	ζ_k
RK-CN $k = 1$	$\frac{4}{15}$	$\frac{8}{15}$	0
RK-CN $k = 2$	$\frac{1}{15}$	$\frac{5}{12}$	$-\frac{17}{60}$
RK-CN $k = 3$	$\frac{1}{6}$	$\frac{3}{4}$	$-\frac{5}{12}$
AB-CN $k = 1, 3$	$\frac{1}{2}$	$\frac{3}{2}$	$-\frac{1}{2}$

Equation (18) is then split into a decoupled set which is a second-order approximation in time to the original equation:

$$\rho^{(4i),n+1} \frac{\hat{u}_i - u_i^n}{\Delta t} = \beta_k (I^{n+1} + I^n) + \gamma_k E^n + \zeta_k E^{n-1} - 2\beta_k \nabla_d P^n \quad (19)$$

$$\rho^{(4i),n+1} \frac{u_i^{n+1} - \hat{u}_i}{\Delta t} = -2\beta_k \nabla_d \delta P^{n+1}. \quad (20)$$

Equation (19) is solved for \hat{u}_i by using the discretizations (10), (14), and (16). Then (20) is divided by $\rho^{(4i),n+1}$ before its discrete divergence is taken to obtain

$$\nabla_d \cdot \left(\frac{1}{\rho^{(4i),n+1}} \nabla_d \delta P \right) = \frac{1}{2\beta_k \Delta t} (\nabla_d \cdot \hat{u}_i - \nabla_d \cdot u_i^{n+1}). \quad (21)$$

A similar Poisson equation with variable coefficient was solved by Bell and Marcus [11] to impose the divergence-free constraint for variable-density flows—see also [12, 13].

Since the transport equation for T has been advanced prior to the momentum equation, the last term in the equation for the pressure variation is known from Eq. (8), written at time $n + 1$. The variable coefficient Poisson equation Eq. (21) for the pressure is solved using the (pre-conditioned) conjugate gradient algorithm. In the case where homogeneous directions exist, it is worth pre-conditioning Eq. (21) by the elliptic operator,

$$\nabla_d \cdot \left(\frac{1}{\langle \rho^{(4i),n+1} \rangle} \nabla_d \right), \quad (22)$$

where $\langle \rangle$ denotes a spatial averaging in the homogeneous directions. In this case one can make use of a FFT-based fast Poisson solver at each iteration of the conjugate gradient algorithm. In the more general case it is worth solving for the modified unknown $\delta P / \sqrt{\bar{\rho}}$ [14]. The advantage of solving Eq. (21) to update the pressure is that the divergence-free constraint is recovered in the inviscid limit, as it has to be from Eq. (8). This is not the case when a backward approximation of $\frac{\partial \rho}{\partial t}$ is used to compute the source term of a linear Poisson equation for δP as proposed earlier [3, 6] (see also Subsection 3.3). The other advantage is that the pressure terms remain energy conserving in the high- R_e number limit, as discussed in the following subsection.

3.2. Fully Conservative Scheme: *DivSC*, *DivSC*_{approx}

Although the previous scheme was found to be accurate, it only conserves momentum and kinetic energy to its own order of accuracy. Experience has shown that the latter quantity

must be conserved exactly if a stable and dissipation-free numerical method is sought. Morinishi *et al.* [9] developed a set of fully conservative (mass, momentum, and kinetic energy) high order schemes for incompressible flow. In the general case of the Navier–Stokes equations without body force, the transport equation for the kinetic energy ρk per unit volume reads

$$\frac{\partial \rho k}{\partial t} + \frac{\partial \rho u_j k}{\partial x_j} = P S_{jj} - \frac{\partial P u_j}{\partial x_j} + \frac{\partial \tau_{ij} u_i}{\partial x_j} - \tau_{ij} S_{ij}. \quad (23)$$

Let us consider a periodic (or infinite) domain so that, after Eq. (23) is integrated over the domain, the flux terms $\partial \rho u_j k / \partial x_j$ and $\partial P u_j / \partial x_j$ make no contribution. Due to the dissipation term $\tau_{ij} S_{ij}$, the question of conservation of the kinetic energy is only relevant in the inviscid limit where $\tau_{ij} = 0$. We know from Eq. (8) that in this limit, the velocity field is divergence-free, that is, $S_{jj} = 0$. Thus global conservation of kinetic energy is a common feature of incompressible and low-Mach number flows. The purpose of this section is to investigate how this property can be extended in discrete space.

Let us define the following discrete approximations of the possible forms for the non-linear term in the momentum equation:

$$\begin{aligned} (Adv.)_i &= \frac{9}{8} \overline{\left(\frac{9}{8} \rho^{(4j)} u_j^{1x_i} - \frac{1}{8} \rho^{(4j)} u_j^{3x_i} \right)} \frac{\delta_1 u_i}{\delta_1 x_j} \\ &\quad - \frac{1}{8} \overline{\left(\frac{9}{8} \rho^{(4j)} u_j^{1x_i} - \frac{1}{8} \rho^{(4j)} u_j^{3x_i} \right)} \frac{\delta_3 u_i}{\delta_3 x_j} \end{aligned} \quad (24)$$

$$\begin{aligned} (Div.)_i &= \frac{9}{8} \frac{\delta_1}{\delta_1 x_j} \left[\overline{\left(\frac{9}{8} \rho^{(4j)} u_j^{1x_i} - \frac{1}{8} \rho^{(4j)} u_j^{3x_i} \right)} \bar{u}_i^{1x_j} \right] \\ &\quad - \frac{1}{8} \frac{\delta_3}{\delta_3 x_j} \left[\overline{\left(\frac{9}{8} \rho^{(4j)} u_j^{1x_i} - \frac{1}{8} \rho^{(4j)} u_j^{3x_i} \right)} \bar{u}_i^{3x_j} \right] \end{aligned} \quad (25)$$

$$(Skew.)_i = \frac{1}{2} ((Adv.)_i + (Div.)_i). \quad (26)$$

The forms $(Adv.)_i$, $(Div.)_i$, and $(Skew.)_i$ are the discrete equivalent to the advective $\rho u_j (\partial u_i / \partial x_j)$, conservative $\partial \rho u_i u_j / \partial x_j$, and skew-symmetric $\frac{1}{2} (\rho u_j (\partial u_i / \partial x_j) + (\partial \rho u_i u_j / \partial x_j))$ forms of the convective term. Note that the discrete operator in Eq. (24) is the same as that in Eq. (10). The following relations holds between these three discrete forms,

$$(Div.)_i = (Adv.)_i + u_i \left(\frac{9}{8} \overline{(Cont.)^{1x_i}} - \frac{1}{8} \overline{(Cont.)^{3x_i}} \right) \quad (27)$$

$$(Skew.)_i = (Adv.)_i + \frac{1}{2} u_i \left(\frac{9}{8} \overline{(Cont.)^{1x_i}} - \frac{1}{8} \overline{(Cont.)^{3x_i}} \right) \quad (28)$$

$$(Skew.)_i = (Div.)_i - \frac{1}{2} u_i \left(\frac{9}{8} \overline{(Cont.)^{1x_i}} - \frac{1}{8} \overline{(Cont.)^{3x_i}} \right), \quad (29)$$

where

$$(Cont.) = \frac{9}{8} \frac{\delta_1 \rho^{(4j)} u_j}{\delta_1 x_j} - \frac{1}{8} \frac{\delta_3 \rho^{(4j)} u_j}{\delta_3 x_j} \quad (30)$$

is the discrete form of the divergence of the momentum vector $\partial \rho u_j / \partial x_j$.

A key assumption in the semi-discrete analysis proposed in Morinishi *et al.* [9] for incompressible flow is that the operator (*Cont.*) is identically zero so that the three forms (*Div.*)_{*i*}, (*Adv.*)_{*i*}, and (*Skew.*)_{*i*} are equivalent. Since (*Div.*)_{*i*} is conservative a priori for the momentum equation and (*Skew.*)_{*i*} is conservative a priori in the kinetic energy equation, a fully conservative scheme is obtained as soon as the velocity constraint $\partial u_j / \partial x_j = 0$ is imposed properly through the pressure correction step. In the present case where the density is not constant, the velocity constraint $\partial u_j / \partial x_j = 0$ (in the high- R_e limit) does not imply that $\partial \rho u_j / \partial x_j$ is zero. Thus the discrete operators (*Div.*)_{*i*}, (*Adv.*)_{*i*}, and (*Skew.*)_{*i*} are not equivalent in the low-Mach number case, meaning that a fully discrete analysis (including the time discretization) must be conducted to achieve conservation of kinetic energy.

A conservative scheme for the momentum can be derived by considering the momentum equation in its divergence form. The first guess for the velocity is obtained by solving

$$\frac{\hat{\rho}^{(4i)} \hat{u}_i - \rho^{(4i),n} u_i^n}{\Delta t} = -\gamma_k (\text{Div.})_i^n - \zeta_k (\text{Div.})_i^{n-1} - 2\beta_k (\text{Pres.})_i^n, \quad (31)$$

where $\hat{\rho}$ can be either ρ^n or ρ^{n+1} or any intermediate value. Then the projection step is written as

$$u_i^{n+1} = \frac{\hat{\rho}^{(4i)}}{\rho^{(4i),n+1}} \hat{u}_i - 2\beta_k \frac{1}{\rho^{(4i),n+1}} \nabla_d \delta P \Delta t, \quad (32)$$

where the Poisson equation for δP must be

$$\nabla_d \cdot \left(\frac{1}{\rho^{(4i),n+1}} \nabla_d \delta P \right) = \frac{1}{2\beta_k \Delta t} \left(\nabla_d \cdot \left(\frac{\hat{\rho}^{(4i)}}{\rho^{(4i),n+1}} \hat{u}_i \right) - \nabla_d \cdot u_i^{n+1} \right). \quad (33)$$

Obviously, Eqs. (31), (32), and (33) constitute a scheme which is momentum conserving. To investigate whether it also conserves kinetic energy, let us multiply Eq. (31) by $\hat{u}_i + u_i^n$ and integrate over the whole domain. The overall contribution of the pressure term involving u_i^n in the kinetic energy equation behaves like

$$\int_V u_i^n (\text{Pres.})_i^n dV = \int_V \left(\frac{9}{8} \overline{u_i^n \frac{\delta_1 P}{\delta_1 x_i}}^{1x_i} - \frac{1}{8} \overline{u_i^n \frac{\delta_3 P}{\delta_3 x_i}}^{3x_i} \right) dV, \quad (34)$$

where

$$\frac{9}{8} \overline{u_i^n \frac{\delta_1 P}{\delta_1 x_i}}^{1x_i} - \frac{1}{8} \overline{u_i^n \frac{\delta_3 P}{\delta_3 x_i}}^{3x_i} = \frac{9}{8} \frac{\delta_1 u_i^n \bar{P}^{1x_i}}{\delta_1 x_i} - \frac{1}{8} \frac{\delta_3 u_i^n \bar{P}^{3x_i}}{\delta_3 x_i} - P \left(\frac{9}{8} \frac{\delta_1 u_i^n}{\delta_1 x_i} - \frac{1}{8} \frac{\delta_3 u_i^n}{\delta_3 x_i} \right). \quad (35)$$

The first two terms of the RHS of Eq. (35) do not contribute because they are in divergence form. The last one is identically zero because the variable coefficient Poisson equation, Eq. (33), is solved with $\nabla_d \cdot u_i^{n+1} = (9/8)(\delta_1 u_i^{n+1} / \delta_1 x_i) - (1/8)(\delta_3 u_i^{n+1} / \delta_3 x_i) = 0$ imposed in the source term (i.e., the projection step Eq. (32) imposes the divergence-free constraint exactly). The other pressure term is $\hat{u}_i (\text{Pres.})_i^n$ and its overall contribution is of order Δt because $\hat{u}_i = u_i^n + O(\Delta t)$ and the integral of $u_i^n (\text{Pres.})_i^n$ is zero.

Using Eq. (26), the overall contribution of the convective terms in the RHS of Eq. (31) may be written as

$$\begin{aligned} & \int_V (\hat{u}_i + u_i^n) (-\gamma_k (Skew.)_i^n - \zeta_k (Skew.)_i^{n-1}) dV \\ & - \gamma_k \int_V \frac{\hat{u}_i + u_i^n}{2} u_i^n \left(\frac{9}{8} (\overline{Cont.})^{n-1x_i} - \frac{1}{8} (\overline{Cont.})^{n-3x_i} \right) dV \\ & - \zeta_k \int_V \frac{\hat{u}_i + u_i^n}{2} u_i^{n-1} \left(\frac{9}{8} (\overline{Cont.})^{n-1x_i} - \frac{1}{8} (\overline{Cont.})^{n-3x_i} \right) dV. \end{aligned} \quad (36)$$

The first integral in Eq. (36) contributes to the order Δt because $(Skew.)_i$ is kinetic energy conserving in nature and because \hat{u}_i , u_i^n , and u_i^{n-1} are equal to the order Δt . Thus an approximation to the order Δt of the overall contribution of the full RHS of Eq. (31) is

$$\begin{aligned} & -\gamma_k \int_V \frac{\hat{u}_i + u_i^n}{2} u_i^n \left(\frac{9}{8} (\overline{Cont.})^{n-1x_i} - \frac{1}{8} (\overline{Cont.})^{n-3x_i} \right) dV \\ & - \zeta_k \int_V \frac{\hat{u}_i + u_i^n}{2} u_i^{n-1} \left(\frac{9}{8} (\overline{Cont.})^{n-1x_i} - \frac{1}{8} (\overline{Cont.})^{n-3x_i} \right) dV. \end{aligned} \quad (37)$$

On the other hand, the contribution of the LHS of Eq. (31) may be written

$$\int_V \frac{\hat{\rho}^{(4i)} (\hat{u}_i)^2 - \rho^{(4i),n} (u_i^n)^2}{\Delta t} dV + \int_V u_i^n \hat{u}_i \frac{\hat{\rho}^{(4i)} - \rho^{(4i),n}}{\Delta t} dV. \quad (38)$$

Comparing Eqs. (37) and (38) it appears that the discrete rate of change of the kinetic energy (the first integral in Eq. (38) is at most of order Δt if one defines the intermediate density as

$$\frac{\hat{\rho} - \rho^n}{\Delta t} = -\gamma_k (Cont.)^n - \zeta_k (Cont.)^{n-1}. \quad (39)$$

In the context of a second order scheme, the same definition of $\hat{\rho}$ was adopted (C. Pierce, private communication) to achieve approximate conservation of kinetic energy.

Multiplying the projection step Eq. (32) by $\hat{u}_i + u_i^{n+1}$ and integrating over the whole domain, the following expression can be derived:

$$\int_V \frac{\rho^{(4i),n+1} (u_i^{n+1})^2 - \hat{\rho}^{(4i)} (\hat{u}_i)^2}{\Delta t} dV = \int_V \hat{u}_i u_i^{n+1} \frac{\hat{\rho}^{(4i)} - \rho^{(4i),n+1}}{\Delta t} dV + O(\Delta t). \quad (40)$$

This shows that the global rate of change of the kinetic energy is of order Δt only if $\hat{\rho}^{(4i)} - \rho^{(4i),n+1} = O(\Delta t^n)$, $n \geq 2$. Unfortunately, n is only 1 in the most general case.

A conservative scheme ($DivSC_{approx}$) is obtained if one accepts that the state equation, Eq. (4), is verified to the order Δt only, viz.,

$$\rho^{n+1} = \hat{\rho} = \frac{P_o}{T_{n+1}} + O(\Delta t). \quad (41)$$

In this case, the error in the global kinetic energy conservation is at most of order Δt for each sub-step. If one accounts for the cancellation of error in the full third-order Runge–Kutta procedure, it can be shown that the error is actually of order Δt^3 . The same order

was achieved for incompressible flows in Morinishi *et al.* [9]. The Adams–Bashforth time integration is not considered here since it is unstable for inviscid flows (see Section 4).

3.3. Scheme with Approximate Poisson Equation: $DivSC_\rho$

The third algorithm is similar to the one in Subsection 3.2 except for the Poisson equation for the pressure. It is briefly described in this section for completeness. One keeps the same scheme, Eq. (31), for the first guess of the velocity. The projection step is now written as

$$\rho^{(4i),n+1} u_i^{n+1} = \hat{\rho}^{(4i)} \hat{u}_i - 2\beta_k \nabla_d \delta P \Delta t, \quad (42)$$

where the Poisson equation for δP must be

$$\nabla_d \cdot \nabla_d \delta P = \frac{1}{2\beta_k \Delta t} (\nabla_d \cdot (\hat{\rho}^{(4i)} \hat{u}_i) - \nabla_d \cdot (\rho^{(4i),n+1} u_i^{n+1})). \quad (43)$$

The divergence of momentum at time level $n + 1$ is unknown but it can be assessed by using the continuity equation:

$$\nabla_d \cdot (\rho^{(4i),n+1} u_i^{n+1}) = -\frac{\delta \rho^{n+1}}{\delta t}. \quad (44)$$

Since the temperature field is advanced first, the equation of state, Eq. (4), can be used to compute the density at level $n + 1$. Then a backward discretization for the time derivative of density can be used to assess the source term in Eq. (43). Since δP is of order Δt , the source term in the Poisson equation is of order Δt too. Since in this source term the time derivative of density is divided by the time step (see Eq. (43)), the backward discretization must be at least second-order accurate for consistency. A possible choice is

$$\frac{\delta \rho^{n+1}}{\delta t} = \frac{[(\Delta t_n + \Delta t_{n-1})^2 - \Delta t_n^2] \rho^{n+1} - (\Delta t_n + \Delta t_{n-1})^2 \rho^n + \Delta t_n^2 \rho^{n-1}}{\Delta t_n \Delta t_{n-1} (\Delta t_n + \Delta t_{n-1})}. \quad (45)$$

Obviously, Eqs. (31), (42), and (43) constitute a scheme which is momentum conserving. However, it is not kinetic energy conserving because the divergence-free constraint is not recovered in the inviscid limit. Thus the overall contribution of the pressure term does not vanish as in Subsection 3.2. The test cases in Section 5 show that this algorithm is less stable and accurate than $DivSC$ and $AdvSC$. However, the Poisson equation, Eq. (43), is with constant coefficient and can be solved very efficiently. Subsequently, it seems that this approach was used in all the previous studies dealing with low-Mach number flows [1–8]. The approach involving a Poisson equation with variable coefficient as in Subsections 3.1 and 3.2 was preferred in [11–13].

The conservative properties of the schemes described in Section 3 are summarized in Table II. Recall that $AdvSC$, $DivSC$, and $DivSC_\rho$ stand for the schemes discussed in Subsections 3.1, 3.2, and 3.3, respectively. $DivSC_{\text{approx}}$ denotes the case where the state equation Eq. (4) is not enforced exactly—see Subsection 3.2. The particular conservation properties of each of the schemes considered are denoted by a cross. The columns “convective,” “pressure,” and “projection” refer to kinetic energy conservation with respect to the convection and pressure terms (in the momentum equation) and the projection step, respectively. The table also specifies whether or not the exact state equation (4) is enforced.

TABLE II
Properties of the Different Algorithms

	Mass	Momentum	Convective	Pressure	Projection	State
<i>AdvSC</i>	x			x		x
<i>DivSC</i>	x	x	x	x		x
<i>DivSC</i> _{approx}	x	x	x	x	x	
<i>DivSC</i> _{ρ}	x	x	x			x

Note. A scheme has a given property if the entry is checked.

From the channel flow computations performed in the course of this study, (see Subsection 5.5), the CPU time required per iteration for *DivSC* is between 1.2 and 2 times higher than that needed for *DivSC* _{ρ} ; the higher the temperature ratio, the higher the cost. The cost difference arises mainly through the resolution of the variable coefficient Poisson equation (33). This difference is expected to decrease in cases with complex geometry which require the use of an iterative method for Eq. (43) of algorithm *DivSC* _{ρ} . Recall that a FFT-based fast Poisson solver has been used in the present study. The cost for algorithm *AdvSC* is slightly smaller than for *DivSC* since the continuity equation is not advanced in time.

4. ERROR ANALYSIS

In order to investigate the stability limit of the previous algorithms in the 1D linear case, a Fourier analysis was performed which is presented in Subsection 4.1. Convergence tests are then described in Subsections 4.2 and 4.3 to show both time and space accuracy of the methods developed in this paper.

4.1. Stability

A Von Neumann analysis of the linear 1D convection-diffusion equation

$$\frac{\partial u}{\partial t} + c \frac{\partial u}{\partial x} = \nu \frac{\partial^2 u}{\partial x^2} \quad (46)$$

was performed to assess the stability of the first step of the algorithms described in the previous section. Note that in this simple case the forms $(Adv)_i$ and $(Div)_i$ in Eqs. (24) and (25) reduce to (convection velocity c omitted)

$$\frac{\partial u}{\partial x} \equiv \frac{27(u_{i+1} - u_{i-1}) - (u_{i+3} - u_{i-3})}{48h}. \quad (47)$$

Also, in the case of constant viscosity, the spatial differencing Eq. (16) for the diffusion term is proportional to

$$\frac{\partial^2 u}{\partial x^2} \equiv \frac{-1460u_i + 783(u_{i+1} + u_{i-1}) - 54(u_{i+2} + u_{i-2}) + (u_{i+3} + u_{i-3})}{576h^2}. \quad (48)$$

Denoting the semi-discrete operator \mathcal{A}_4 such that $\partial u / \partial t = \mathcal{A}_4 u_i^n$, it is straightforward to demonstrate using Eqs. (47) and (48) that its Fourier transform $\hat{\mathcal{A}}_4(\omega)$ may be expressed as

$$\Delta t \hat{\mathcal{A}}_4 = -j \frac{CFL}{24} (27 \sin(\omega h) - \sin(3\omega h)) + \frac{Fo}{288} (-730 + 783 \cos(\omega h) - 54 \cos(2\omega h) + \cos(3\omega h)), \quad (49)$$

where $CFL = c \Delta t / h$ and $Fo = v \Delta t / h^2$ denote the Courant and the Fourier numbers respectively while $j^2 = -1$. For the full discretization to be stable, it is necessary that $\Delta t \hat{\mathcal{A}}_4(\omega)$ be contained in the stability region of the time differencing for all values of ω . The locus of $\Delta t \hat{\mathcal{A}}_4(\omega)$ is shown for different values of CFL and Fo together with the stability region of the Adams–Bashforth algorithm in Fig. 1b. Figure 1a shows the same quantities for the classical second-order semi-discretization where $\Delta t \hat{\mathcal{A}}_2 = -j CFL \sin(\omega h) + 2Fo(\cos(\omega h) - 1)$. As expected, the fourth-order scheme appears to be less stable than the second-order one. Since both semi-discretizations $\hat{\mathcal{A}}_2$ and $\hat{\mathcal{A}}_4$ are anti-symmetric in the convection part, the Adams–Bashforth time differencing is always unstable for $Fo = 0$. For finite viscosity, both spatial schemes are stable, although the domain of stability is smaller for the present formulation. For pure diffusion, viz. $CFL = 0$, the critical Fourier number for $\hat{\mathcal{A}}_2$ is 0.25 whereas for $\hat{\mathcal{A}}_4$ it is 0.197. The same analysis was performed for the third-order Runge–Kunta time differencing (not shown). In the absence of viscosity ($Fo = 0$), the critical CFL number is 1.54, compared to the equivalent value for the classical second-order scheme which is 1.8.

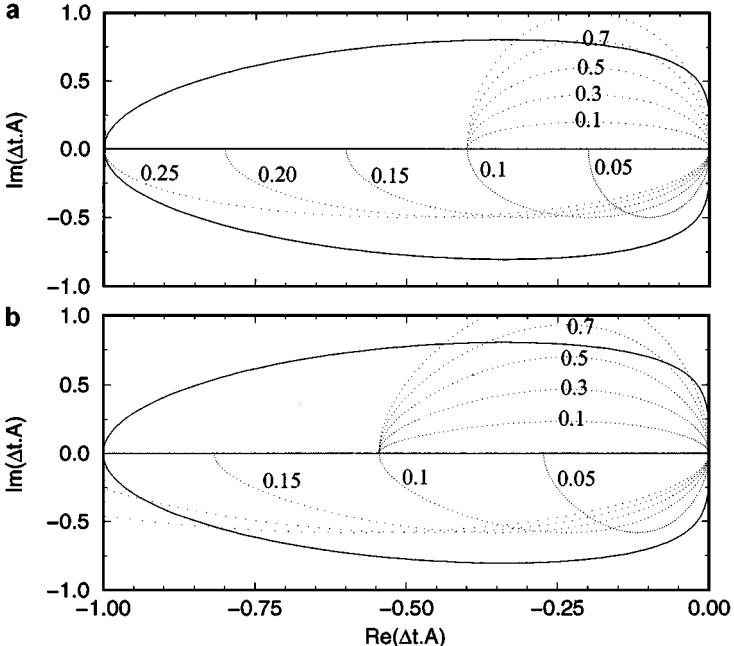


FIG. 1. Von Neumann stability analysis: (a) second-order centered scheme, (b) present scheme. —, Adams–Bashforth stability region; . . . , locus of $\Delta t \hat{\mathcal{A}}$. Upper half-planes, $Fo = 0.1$ and CFL varies (increment 0.2); bottom half-planes, $CFL = 0.5$ and Fo varies (increment 0.05).

4.2. Time Accuracy

Time accuracy of the numerical methods described in Section 3 was assessed through a numerical experiment in the 2D periodic domain $[0, 2\pi] \times [0, 2\pi]$. The initial condition was set to $u(x, y, t = 0) = \partial\psi/\partial y + \sin(x) \times \cos(y)/(RePr)$, $v(x, y, t = 0) = -\partial\psi/\partial x + \cos(x) \times \sin(y)/(RePr)$, $P(x, y, t = 0) = 0$ and $T(x, y, t = 0) = 1 - \cos(x) \times \cos(y)$ so that the constraint on the velocity field, Eq. (8), is satisfied initially (assuming constant viscosity). The stream function is equal to $\psi = \cos(x) \times \cos(y)$. Several simulations were performed with $Re = 2000$ and $Pr = 1$ and using a uniform 16×16 grid. The number of time-steps (Adams–Bashforth time integration) used to cover a given time interval was increased from 4 to 32, corresponding to Courant numbers in the range $[0.011, 0.088]$. One more computation was performed for which 256 iterations were necessary in order to generate a highly accurate reference solution. The error in the solutions obtained at larger Δt was computed as the \mathcal{L}_2 -norm of the difference in the solution for a given time increment when compared with the reference solution. Regarding the u -velocity, second-order accuracy is demonstrated in Fig. 2 for all the considered algorithms. The same convergence rate was obtained for the v -velocity and for the temperature. An interesting point to note is that the convergence rate for the pressure is only unity. This is not related to the density variations considered here since the same phenomenon is true for the incompressible Morinishi’s algorithm [9]. As discussed by Perot [15], the order of accuracy of the pressure update does not affect the order of accuracy of the velocity field.

4.3. Spatial Accuracy

Other numerical experiments were conducted to study the spatial accuracy of the algorithms. Only the results for *AdvSC* and *DicSC* are discussed in this section since the spatial differencing for all the previous algorithms reduces to one or the other of these two approaches. The initial condition was the same as that used in Subsection 4.2 and four different uniform grids were considered, namely ones containing 16^2 , 48^2 , 144^2 , and 432^2 cells. The solutions obtained after a given integration time are denoted S_{16} , S_{48} , S_{144} , and S_{432} respectively (S may denote either one of the velocity components, the pressure, or the

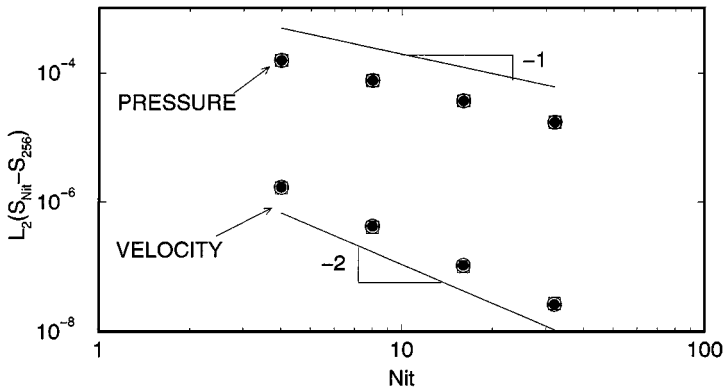


FIG. 2. \mathcal{L}_2 -norm of the difference between the solution (u -velocity or pressure) and the “exact” solution as a function of the number of iterations performed to compute the given time interval. \circ , *DivSC*; \bullet , *DicSC*_{approx}; \square , *AdvSC*.

TABLE III
Spatial Convergence Rates for *AdvSC* and *DivSC*

	<i>u</i>	<i>v</i>	<i>P</i>	<i>T</i>
<i>AdvSC</i>	3.89; 4.19	3.89; 4.19	4.06; 4.35	3.87; 3.92
	3.89; 4.18	3.89; 4.18	4.04; 4.36	3.87; 3.90
<i>DivSC</i>	3.88; 3.60	3.88; 3.60	4.06; 3.58	3.87; 3.99
	3.87; 3.59	3.87; 3.59	4.03; 3.59	3.87; 3.98

Note. For each entry, the first two values are determined from Eq. (50) with $(n, m) = (16, 48)$ and $(n, m) = (48, 144)$ while the third and fourth values are from Eq. (51) with $(n, m, k) = (16, 48, 144)$ and $(n, m, k) = (48, 144, 432)$.

temperature). This grid sampling was chosen so that all the velocity and pressure nodes of the 16^2 staggered grid are common to all the grids. The \mathcal{L}_2 -norm of the difference between two solutions can be computed on the coarser mesh without further interpolation. It is believed that in doing so one obtains a more reliable estimation of the spatial convergence rate. For all the computations, the Courant number was kept to a small value ($CFL \approx 5 \times 10^{-3}$) so that the errors associated with the time differencing are small with respect to the spatial differentiation. The number of iterations was 16 for the coarsest grid, 432 for the finest one. The convergence rate σ was computed in two different ways. First the solution obtained with the finest grid (\mathcal{S}_{432}) was interpreted as the “exact” solution. Then assuming a spatial error of the form $C \cdot h^\sigma$, the convergence rate was estimated by

$$\sigma \approx \frac{\ln[\mathcal{L}_2(\mathcal{S}_n - \mathcal{S}_{432})] - \ln[\mathcal{L}_2(\mathcal{S}_m - \mathcal{S}_{432})]}{\ln 3}, \quad (50)$$

where (n, m) is either $(16, 48)$ or $(48, 144)$. The other way of computing σ does not require knowledge of the exact solution [11] and assumes that the \mathcal{L}_2 -norm of the difference between two successive grid levels is proportional to the grid spacing in the coarser mesh. It is then straightforward to show that

$$\sigma \approx \frac{\ln[\mathcal{L}_2(\mathcal{S}_n - \mathcal{S}_m)] - \ln[\mathcal{L}_2(\mathcal{S}_m - \mathcal{S}_k)]}{\ln 3}, \quad (51)$$

where (n, m, k) is either $(16, 48, 144)$ or $(48, 144, 432)$. Both estimates Eqs. (50) and (51) were applied to the velocity components, the pressure, and the temperature fields. The results are shown in Table III. The fourth-order accuracy of the methods presented in Section 3 is demonstrated.

5. NUMERICAL TESTS

A few test cases were designed to illustrate the performance of the numerical methods described in the previous section, including the computation of a 1D high-amplitude pulse of density in an inviscid flow, a 1D small amplitude pulse of density in a viscous flow, 2D random fluctuations of velocity and temperature, 2D small amplitude fluctuations in a channel, and 3D large amplitude fluctuations in a channel. A brief description of the

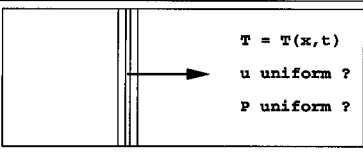
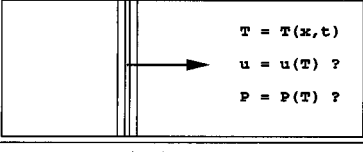
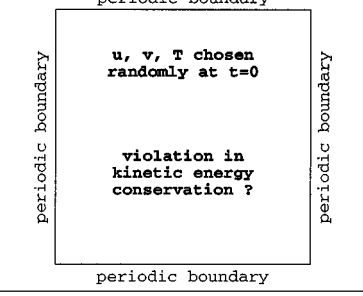
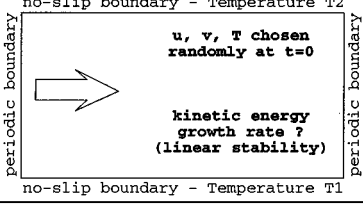
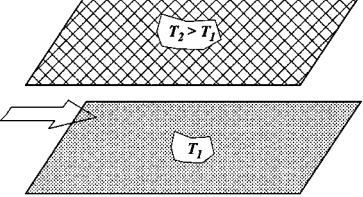
Description	Sketch	Objectives
<p>section 5.1</p> <ul style="list-style-type: none"> - 1D - inviscid - large fluctuations 		<ul style="list-style-type: none"> - constant pressure
<p>section 5.2</p> <ul style="list-style-type: none"> - 1D - viscous - small fluctuations 		<ul style="list-style-type: none"> - 1D wave structure - Prandtl number effect
<p>section 5.3</p> <ul style="list-style-type: none"> - 2D - inviscid - non-linear 		<ul style="list-style-type: none"> - kinetic energy conservation
<p>section 5.4</p> <ul style="list-style-type: none"> - 2D - viscous - linear 		<ul style="list-style-type: none"> - most unstable mode selection - modal structure - growth rate
<p>section 5.5</p> <ul style="list-style-type: none"> - 3D - viscous - non-linear 		<ul style="list-style-type: none"> - DNS capability - higher density ratios on coarse grid

FIG. 3. Basic test cases for the low-Mach number algorithms.

different test cases is provided in Fig. 3 which also summarizes the assumptions and the objectives for each test case. The time integration must be RK-CN for the inviscid test cases while either RK-CN or AB-CN can be used for the viscous ones. RK-CN has been used for the test cases in Subsections 5.1, 5.3, and 5.4. The time integration is AB-CN in Subsections 5.2 and 5.5. The results obtained are discussed in more detail in the following subsections.

5.1. The 1D Euler Convection

If the Peclet number is infinite, the velocity field must be divergence-free, that is, u must be constant in 1D. Also, the pressure should remain constant. To test the ability of the different formulations to reproduce this feature of Eqs. (1)–(3), consider the domain $0 \leq x \leq 1$, periodic in x . The initial condition is $u = u_0 = 1$, $P = 0$, and $T = 1 + A \exp[-((x - x_0)/a)^2]$

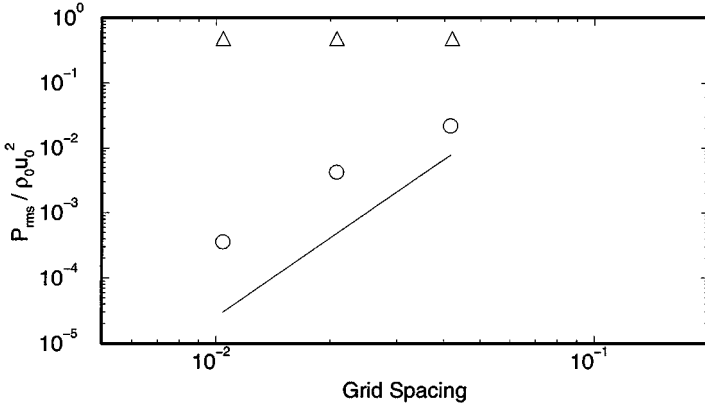


FIG. 4. Root-mean-square of pressure as a function of the grid spacing at time $t = 20a/u_0$. —, $P_{\text{rms}} \propto \Delta x^4$; ●, *DivSC*; □, *AdvSC*; △, *DivSC_ρ*.

with $A = 1$, $x_0 = 0.5$, and $a = 0.05$. When the grid contains $N_x = 24$ points, only 6 points are used to describe a Gaussian perturbation. Figure 4 shows $P_{\text{rms}}/\rho_0 u_0^2$ as a function of the grid spacing (this quantity is also the error in P_{rms} since the exact solution is $P_{\text{rms}} = 0$). Three grid levels were considered: 24, 48, and 96 points in x . The rms of pressure is assessed for the time $t = 20a/u_0$. The CFL number is of order 0.5 in all cases. Since *AdvSC* renders the exact solution for this particular test case, the corresponding results are not included in the figure. The divergence is zero for *DivSC* because it is explicitly enforced through the Poisson equation, Eq. (33). On the other hand, the divergence-free constraint is not enforced in the algorithm *DivSC_ρ* and the velocity does not remain constant. Since the error in the backward difference equation, Eq. (45), is proportional to a high-order time derivative of density, one expects the error in the source term of the Poisson equation, Eq. (43), to increase with the amplitude of the density perturbation. In other words, the error in the divergence of velocity should be proportional to the amplitude A of the initial temperature perturbation, at least for small values of A . This is illustrated in Fig. 5.

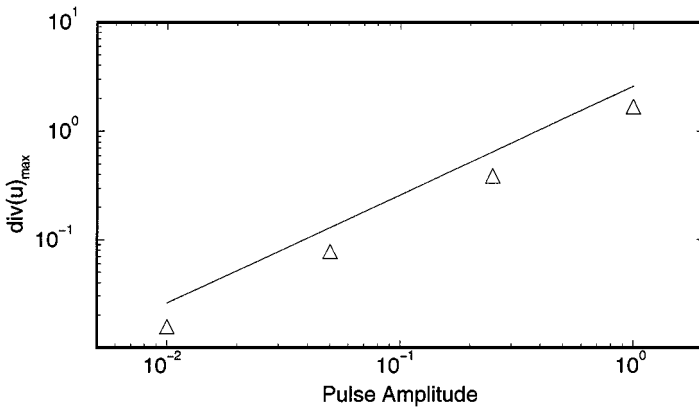


FIG. 5. Maximum of divergence $\frac{du}{dx}$ as a function of the amplitude A of the temperature perturbation. —, $\frac{du}{dx} \propto A$; △, *DivSC_ρ*.

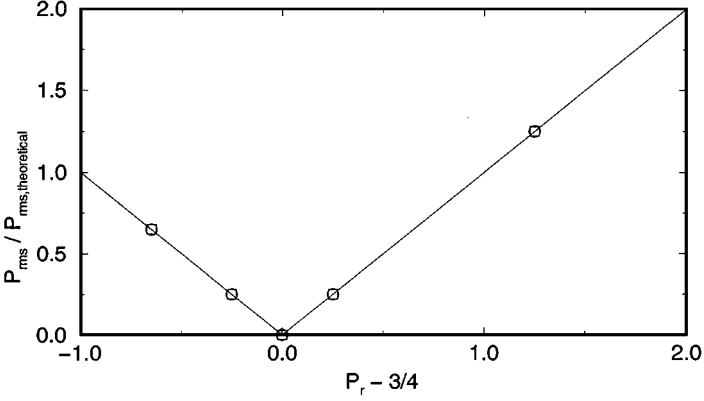


FIG. 6. Root-mean-square of pressure as a function of the Prandtl number. P_{rms} is non-dimensionalized by its theoretical value as a function of T' , viz. $(4/3R_e^2 P_r^2)(\partial^2 T'/\partial x^2)_{rms}$. Note that the abscissa is $P_r - 3/4$. —, exact solution Eq. (54); \circ , *DivSC*; \square , *AdvSC*.

5.2. Small 1D Perturbations

In the case where the Reynolds number is finite but where the perturbation in temperature is small ($A \ll 1$), the analytical resolution of Eqs. (1)–(3) can be conducted and the structure of the perturbation which propagates is given by

$$\rho' = -\frac{\rho_0}{T_0} T' \quad (52)$$

$$u' = \frac{1}{R_e P_r} \frac{\partial T'}{\partial x} \quad (53)$$

$$p' = \frac{4}{3R_e^2 P_r^2} \frac{\partial^2 T'}{\partial x^2} \left(P_r - \frac{3}{4} \right). \quad (54)$$

An interesting feature is that the pressure fluctuation should vanish in the limit $P_r = \frac{3}{4}$. Figure 6 shows the error in Eq. (54) in the case $N_x = 24$, $a = 0.05$, $A = 0.01$, and $R_e = 50$. The initial condition is uniform for u and P and the physical time simulated is large enough ($t \approx 160a/u_0$) so that the values reported in the figure are asymptotic values. Both *DivSC* and *AdvSC* are in complete agreement with the theory. In order to derive the analytical relationship Eq. (54) between p' and T' , one makes use of the constraint on the divergence of the velocity field which reduces to Eq. (53) in the present case. Since this constraint is not properly imposed by the algorithm *DivSC*, there is no reason to expect that this method gives results in agreement with Eq. (54). Instead, the pressure fluctuation should only depend on the initial temperature fluctuation (shape and amplitude). Accordingly, Fig. 7 shows that the rms of pressure behaves like P_r^2 when scaled by $(4/3R_e^2 P_r^2)(\partial^2 T'/\partial x^2)_{rms}$ as in Fig. 6.

5.3. The 2D Random Perturbations

To validate the results of Subsection 3.2 with numerical tests, inviscid flow simulations were performed on a 2D periodic domain. The analytical solution dictates that the total momentum in each direction $\langle \rho u_i \rangle$ and total kinetic energy $\langle K \rangle = \frac{1}{2} \langle \rho u_i^2 \rangle$ should be

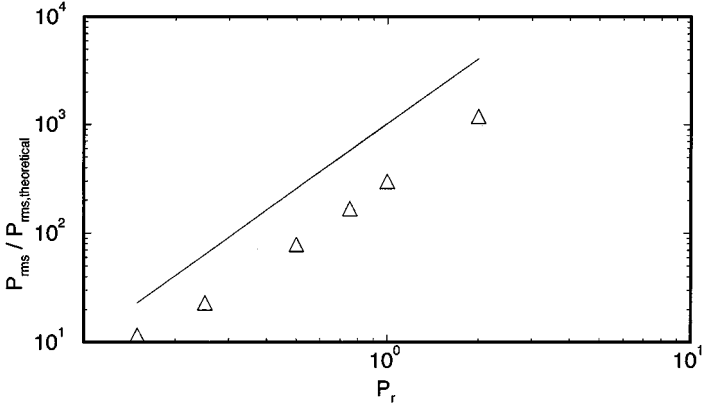


FIG. 7. Root-mean-square of pressure as a function of the Prandtl number. P_{rms} is non-dimensionalized by its theoretical value as a function of T' , viz. $(4/3R_e^2 P_r^2)(\partial^2 T'/\partial x^2)_{\text{rms}}$. The abscissa is P_r . —, $P_{\text{rms}} \propto P_r^2$; Δ , DivSC_ρ .

conserved in time. The domain is $0 \leq x \leq L$, $0 \leq y \leq L$ and a 24×24 mesh is used. Solenoidal velocity fields are used as the initial condition together with random temperature fluctuations. The initial mean kinetic energy is of order 1.5 while $T_{\text{rms}} \approx 0.15\langle T \rangle$ at $t = 0$. Figure 8 shows the relative error for the total kinetic energy $\langle K_0 - K \rangle / \langle K_0 \rangle$ after an integration time of $t = 0.125L / \sqrt{\langle K_0 \rangle}$. As expected from Subsection 3.2, the error for the scheme DivSC does not behave like Δt^3 because of the violation in the conservation of kinetic energy related to the non-linear equation of state Eq. (4). On the other hand, it appears that the same scheme with the approximate equation of state (Eq. (41)) conserves the global kinetic energy to the order Δt^3 as predicted in Subsection 3.2. The scheme DivSC_ρ violates the conservation of kinetic energy even if the approximate equation of state is used. This is because the divergence-free constraint is not recovered in the inviscid limit when the Poisson equation, Eq. (43), is used. Thus the pressure term is not kinetic energy conserving for DivSC_ρ (see Subsection 3.2, Eq. (35)) which was found unstable for the present test case. The convective terms in the algorithm AdvSC violate the conservation of kinetic energy.

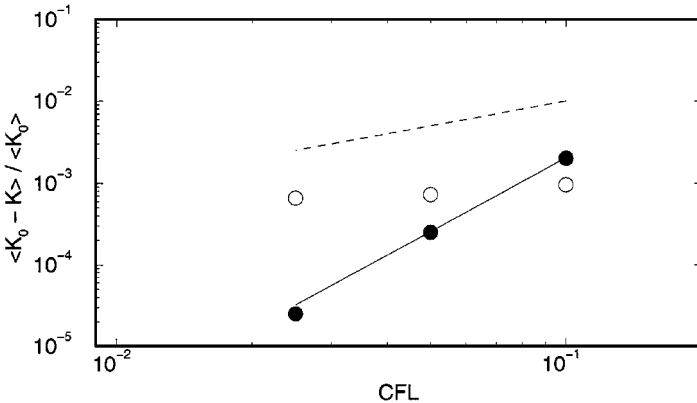


FIG. 8. Kinetic energy conservation error ($\langle K_0 - K \rangle / \langle K_0 \rangle$) as a function of the Courant number. —, Δt^3 behavior; ---, Δt behavior; \circ , DivSC ; \bullet , $\text{DivSC}_{\text{approx}}$.

5.4. Linear Stability in a Channel

To check the accuracy of the algorithm in the case where the physical properties vary in space and time through the temperature, the evolution of low amplitude eigenmodes in laminar channel flow is simulated. The linear stability problem in a channel flow between two isothermal walls with temperature $T_1 = 1 - \frac{\delta T}{2}$ and $T_2 = 1 + \frac{\delta T}{2}$ was studied by Suslov and Paolucci [16] under the low-Mach number assumption. They found that the critical Reynolds number increases with the parameter $\frac{\delta T}{2}$. It is of order 40,000 for $\frac{\delta T}{2} = 0.4$, compared to 5,772 in the isothermal case ($\delta T = 0$). In their analysis the dimensionless thermal conductivity and dynamic viscosity are given by Sutherland's law,

$$k(T) = T^{3/2} \frac{1 + S_k}{T + S_k}, \quad \mu(T) = T^{3/2} \frac{1 + S_\mu}{T + S_\mu}, \quad (55)$$

where $S_k = 0.648$ and $S_\mu = 0.368$ for air at $T_{\text{ref}} = 300\text{K}$ and normal pressure. The molecular Prandtl number is 0.76. In the computation, the length of the periodic domain in x is $L = 2\pi/\alpha$, where α is the wave number of the mode of interest. The initial condition consists of a small amplitude (0.01%) random noise on u, v superimposed to the laminar solution of the problem (Suslov and Paolucci [16]). A stretched grid is used in the normal direction in order to capture the eigenvector accurately near the walls. The wall normal velocity points are distributed according to a hyperbolic tangent function ($j = 0, 1, 2, \dots, N$),

$$y_v(j) = y_{j+\frac{1}{2}} = \frac{\tanh(\gamma(2j/N - 1))}{\tanh(\gamma)}. \quad (56)$$

A typical result is shown in Fig. 9. In this case the resolution is 24×100 with $\gamma = 2$ for the stretching parameter. The CFL number is fixed at 1. The length of the domain is $L = 2.4\pi h$ ($\alpha = \frac{5}{6} \frac{1}{h}$) and the Reynolds number is 45,000, based on the maximum velocity and the channel half-height h . The temperature ratio is $T_2/T_1 = 2.33$, i.e., $\frac{\delta T}{2} = 0.4$. For these conditions, the flow is linearly unstable (see Suslov and Paolucci [16]). The code predicts a reasonable (5% error) energy growth rate even if the number of grid points is rather small in the direction normal to the wall. Note that a fairly long time ($10h/u_\tau$) is

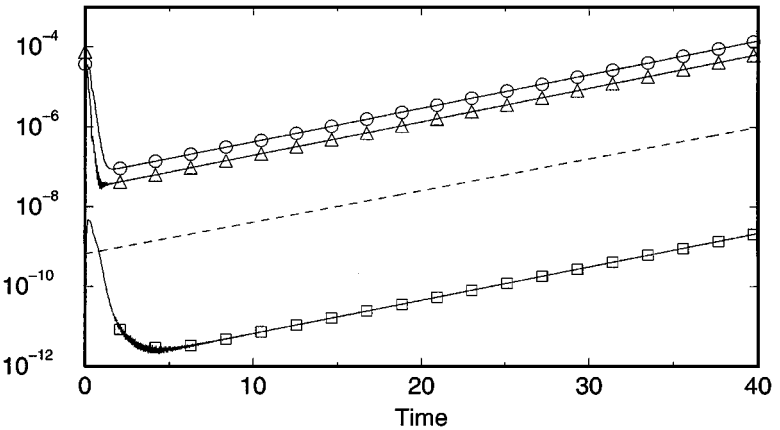


FIG. 9. Time evolution of the global energy of the fluctuations in the computational domain. ----, linear stability theory (Suslov and Paolucci); \circ -, $\langle u^2 \rangle$; \triangle -, $\langle v^2 \rangle$; \square -, $\langle T^2 \rangle$. Unit of time is h/u_τ .

needed for the mode to settle in. Once the transition phase is finished, the temperature and the two velocity components develop with exactly the same rate, as dictated by linear stability theory. The four schemes $AdvSC$, $DivSC$, $DivSC_{approx}$, and $DivSC_\rho$ give very similar results ($AdvSC$ shown). The success of $DivSC_\rho$ in this test case is due to the fact that only small temperature (density) perturbations are considered for comparison with the linear theory. Thus the error in the time derivative of density is negligible compared to the other space discretization errors (see Subsection 5.1) and the algorithm provides a reasonable answer.

The algorithms $AdvSC$ and $DivSC$ have been found to give consistent results over a wide range of flow and numerical conditions (temperature ratio from 1.01 to 2.33, grid size from 16×100 to 56×350). In order to establish the qualitative nature of the algorithms, typical 2D plots of velocity, pressure, and temperature fluctuations are shown in Fig. 10 in which the complete domain is shown. Figure 11 shows the iso-lines for u and T for the region near the cold (bottom) wall of the channel. The conditions are the same as for Fig. 9. The modal

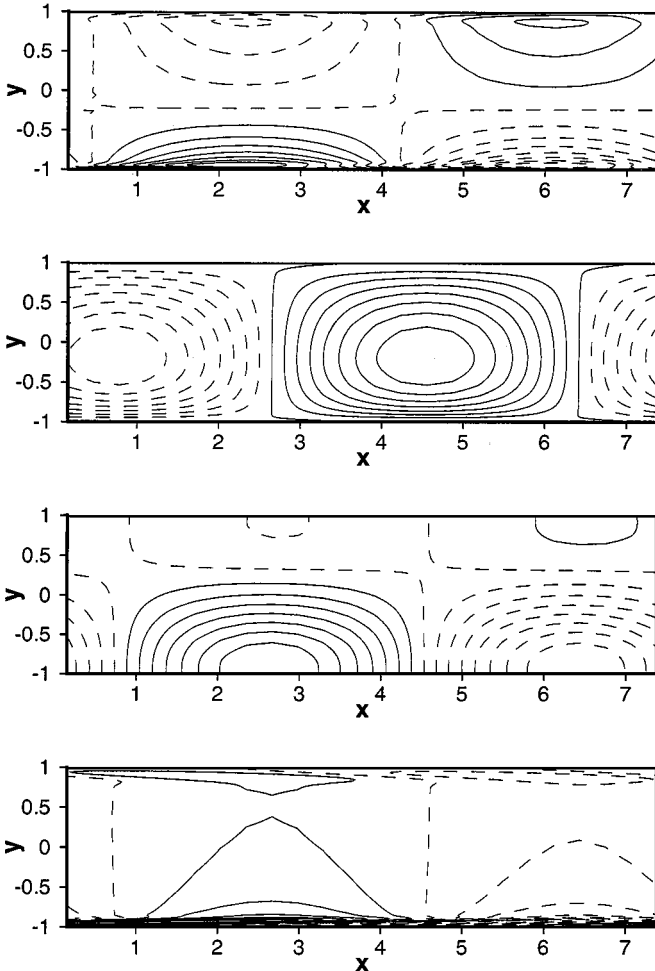


FIG. 10. Fluctuations associated with the most unstable linear mode for $\alpha = \frac{5}{6} \frac{1}{h}$, $Re = 45,000$, and $\frac{\delta T}{2} = 0.4$. From top to bottom, u -velocity, v -velocity, pressure, and temperature fluctuations. Dashed iso-lines are negative values. The mean flow is from left to right and the bottom wall is the cold one. The mesh size is 24×100 .

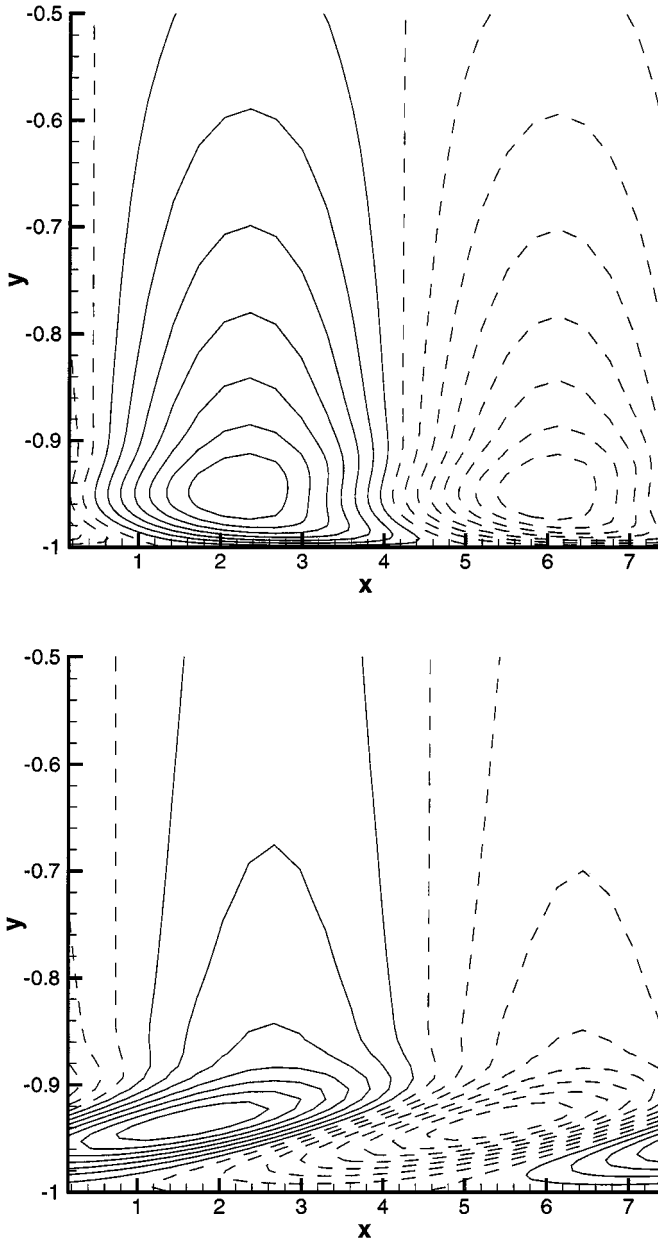


FIG. 11. Fluctuations of u -velocity (top) and temperature (bottom). See the previous figure for the conditions. Zoom of region near the cold wall.

structure shown in Figs. 10 and 11 is in very good agreement with previous theoretical findings [17].

5.5. Turbulent Flow in a Channel

Several computations of a turbulent flow in a channel with isothermal walls have been performed. Results from DNSs are presented first. Results from coarse grid computations are then discussed in terms of stability.

5.5.1. Direct numerical simulations. Three direct numerical simulations of a channel flow between two isothermal walls with temperatures T_1 and T_2 were performed. A few results are shown in this section to illustrate the performance of the method proposed in Subsections 3.1 and 3.2. A more detailed analysis can be found in [18, 19]. The computations presented here were performed with *AdvSC*. Details of the test cases adopted are given in Table IV.

The indices 1 and 2 denote the cold and hot wall, respectively. The bulk Reynolds number R_b is based on the bulk velocity and the values of density and dynamic viscosity corresponding to the bulk temperature. The friction coefficients C_f are based on the mean density in the channel and the maximum velocity while the heat flux parameter is defined as $B_q = q_w / \rho_w C_p u_\tau T_w$, where u_τ is the friction velocity $\sqrt{\tau_w \rho_w}$, q_w the heat flux, C_p the constant-pressure specific heat, and T_w the temperature at the wall considered. The friction Reynolds number R_τ is based on the friction velocity. In Case A the temperature is almost uniform and the results may be compared to a previous incompressible DNS performed by Kim *et al.* [20, 21]. In Cases B–C one expects the temperature (density) variations to be large enough to modify the momentum balance through both viscous and inviscid effects. In each case the domain size is $(4\pi h, 2h, 4\pi h/3)$ and the grid contains $120 \times 100 \times 120$ cells. The mean flow is along the x -direction. The statistics were obtained over a time period of order $10h/\bar{u}_\tau$, where $\bar{u}_\tau = (u_{\tau 1} + u_{\tau 2})/2$ is the mean friction velocity. The wall normal velocity points are distributed according to a hyperbolic tangent function, Eq. (56). The grid spacing is equivalent for all cases with $\Delta x^+ \approx 20$, $\Delta y_{\text{wall}}^+ \approx 0.3$, $\Delta y_{\text{max}}^+ \approx 9$, and $\Delta z^+ \approx 6$, where the superscript $+$ denotes the usual wall scaling $U^+ = U/u_\tau$ and $y^+ = \rho_w u_\tau y / \mu_w$. The molecular Prandtl number is 0.76 and the dimensionless thermal conductivity and dynamic viscosity are chosen to be proportional to $1/\sqrt{T}$ so that the Reynolds number near the hot wall is not (too) small in comparison with its value near the cold wall [18]—see Table IV. It was shown in [19] that the ratio G_r/R_τ^2 (G_r is the Grashof number) is small compared to unity as long as the characteristic length scale of the channel is of order 1 cm or less. This would be a reasonable range for a true experiment with temperature differences typical of a laboratory combustion chamber. Consequently, the buoyancy effects are neglected in the present test cases.

Many experimental data support the validity of the Van Driest [22] transformation for wall-bounded turbulent flows with variable density. This transformation reads

$$U_{VD}^+ = \int_0^{u^+} \left(\frac{\rho}{\rho_w} \right)^{1/2} du^+ = \frac{1}{\kappa} \ln y^+ + C. \quad (57)$$

Figure 12 shows that the present calculations also give support to the transformation (57). When the classical incompressible coordinates are used instead, a logarithmic region can hardly be observed. Note also that Case A is in very good agreement with the law-of-the

TABLE IV
Description of the DNSs

Case	T_2/T_1	$R_{\tau 1} - R_{\tau 2}$	R_b	$10^3 C_{f1}$	$10^3 C_{f2}$	B_{q1}	B_{q2}
A	1.01	182 – 185	2855	6.1	6.1	≈ 0	≈ 0
B	2	195 – 164	2810	7.0	5.0	-0.018	+0.016
C	4	211 – 151	2818	8.2	4.2	-0.041	+0.029

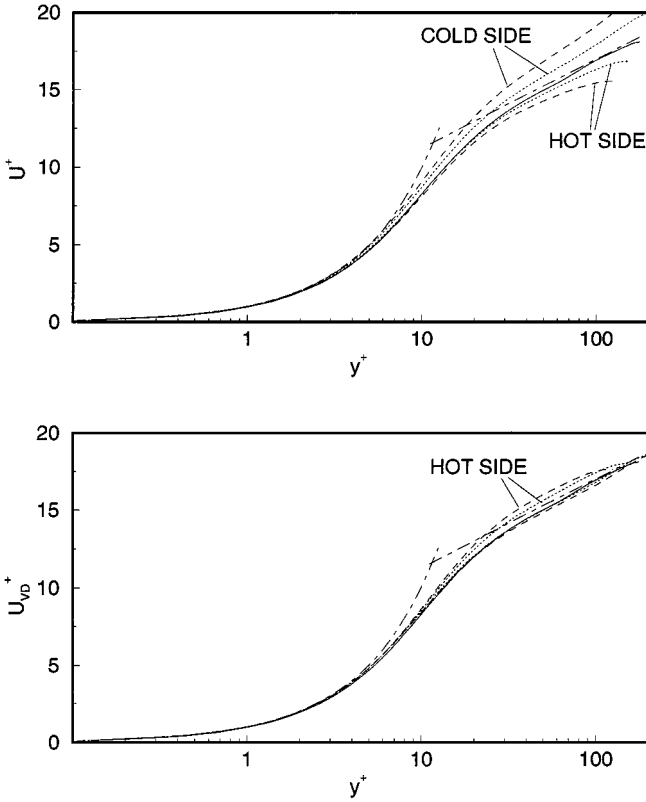


FIG. 12. Mean velocity profiles. Top figure without the Van Driest transformation Eq. (57); bottom figure with it. —, Case A; ···, Case B; ---, Case C; — · —, $u^+ = y^+$ and $u^+ = 2.5 \ln y^+ + 5.5$.

wall $u^+ = 2.5 \ln y^+ + 5.5$, as proposed in [20]. Regarding the velocity fluctuations and the Reynolds stress $-\overline{u'v'}$, this computation is also in very good agreement with the incompressible DNS data from Kim *et al.* [20]—see Fig. 13. Figure 14 shows a very good qualitative agreement for the temperature fluctuations in Case A when compared to a passive scalar

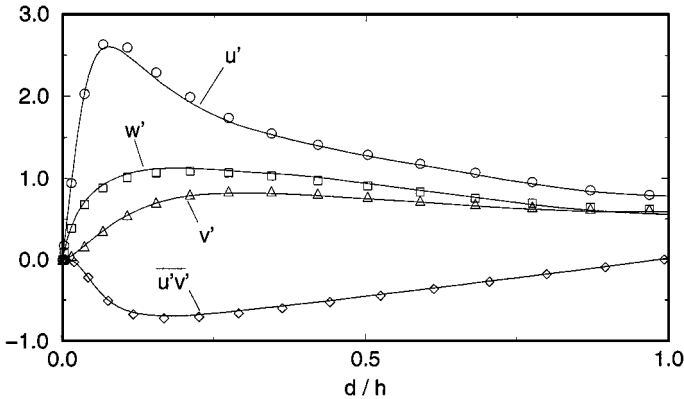


FIG. 13. Root-mean-square velocities and Reynolds stress from the DNS (Case A) in wall units. Symbols from Kim *et al.* [20]. \circ , u'_{rms} ; \triangle , v'_{rms} ; \square , w'_{rms} ; \diamond , $\overline{u'v'}$.

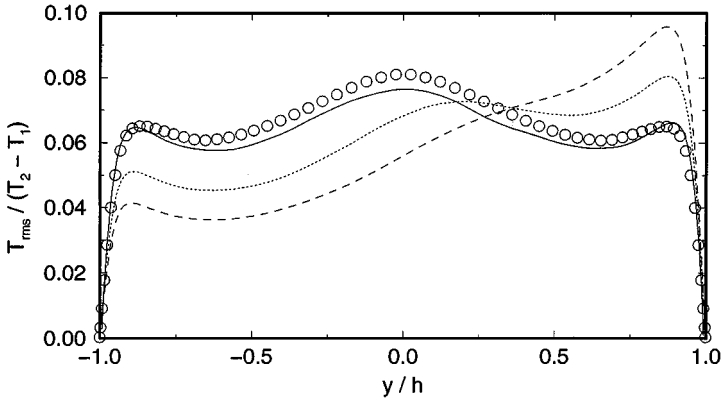


FIG. 14. Root-mean-square of the temperature fluctuations from Cases A, B, and C, non-dimensionalized by the temperature difference $T_2 - T_1$. —, Case A; ···, Case B; ---, Case C; ○, passive scalar case with $Pr = 0.71$, Kim and Moin [21].

calculation [21]. The slight quantitative difference is most likely due to the difference in Prandtl numbers ($Pr = 0.76$ for Case A, $Pr = 0.71$ in [21]). In each of the Cases A–C, the streamwise velocity-temperature correlation coefficient is very high near the walls (of order 0.95, as in [21]). Subsequently the temperature field reveals the elongated streaky structure of turbulence near the wall. This is illustrated for Case A in Fig. 15 which shows iso-lines of u -velocity and temperature fluctuations in a plane located approximately 5 wall units above the cold wall. As already observed in previous studies, the temperature field is a good

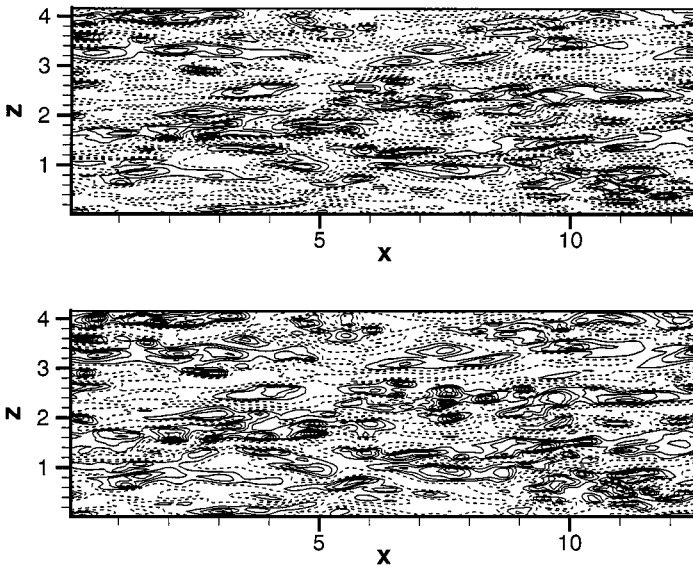


FIG. 15. DNS of a channel flow with $T_2/T_1 = 1.01$ (Case A). The position of $x - z$ plane is $y/h \approx -0.975$, i.e., 5 wall units above the cold wall. The flow is from left to right. Top, contours of u^+ -velocity fluctuations: the range is $[-2, +6]$, the increment is 1, dashed contours represent negative values. Bottom, contours of temperature fluctuations (non-dimensionalized by $T_2 - T_1$): the range is $[-0.04, 0.04]$, the increment is 0.02, dashed contours represent negative values.

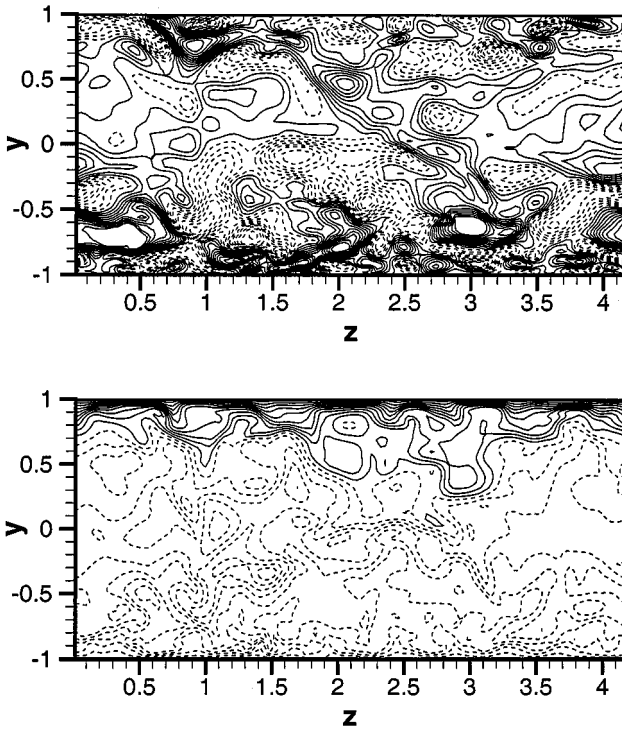


FIG. 16. DNS of a channel flow with $T_2/T_1 = 4$ (Case C). Typical $z - y$ plane. Top, contours of w^+ -velocity: the range is $[-4, +4]$, the increment is 0.4, dashed contours represent negative values. Bottom, contours of temperature (non-dimensionalized by T_1): the range is $[1, 4]$, the increment is 0.15, dashed contours represent values smaller than $(T_2 + T_1)/2$.

marker for the bursting events, as shown in Fig. 16 for Case E. Of interest also in this figure is the position of the iso-line $T = (T_1 + T_2)/2$. It is clearly skewed towards the hot wall. Indeed, the Reynolds number in the hot region being smaller than in the cold region—see Table IV—the linear sub-layer close to the top wall is thicker. As a consequence, the mean temperature decreases faster near the top wall ($y = +1$) than it increases near the bottom one ($y = -1$). The iso-lines of the w -velocity in Fig. 16 reveal the streamwise vortices responsible for the streaky structure shown in Fig. 15.

5.5.2. Coarse grid computations. In algorithms of the type $DivSC_\rho$ where a constant coefficient Poisson equation (43) is solved for the pressure field, the projection step was found to be the most destabilizing part of the method and density ratios larger than 3 are difficult to compute [6]. In the present algorithms $AdvSC$, $DivSC$, or $DivSC_{approx}$, the exact constraint on the velocity field is enforced. Further computations of the channel flow were performed to see whether the methods proposed are more stable. The numerical and physical parameters are the same as before, except for the grid which is now $32 \times 80 \times 32$. This mesh is too coarse to reproduce all the scales that are dynamically important. However, its resolution makes it computationally inexpensive to use and it was therefore considered suitable for investigating the stability characteristics of the different algorithms. The temperature ratios considered were 1.01, 2, 4, 6, 8, and 10. The algorithm $DivSC_\rho$ is found to be unstable for T_2/T_1 greater than 4. However, the other algorithms $AdvSC$, $DivSC$, and $DivSC_{approx}$ remain stable up to $T_2/T_1 = 10$. Larger temperature gradients were not considered since most of the

low-Mach number flows have temperature ratio less than 10 (piston engine, rocket engines, burners, etc.). For cases $T_2/T_1 \geq 6$, the temperature field was clipped to values greater than the temperature imposed at the cold wall T_1 . Indeed, with this coarse grid, a large amount of energy is present at the highest wave numbers (a common feature with more realistic LESs). For these scales the dispersion phenomenon is important and may generate negative values of temperature if a non-positive scheme is used. The clipping to values greater than the cold wall temperature is justified since there is no physical mechanism which can decrease the temperature below this point. It emulates a TVD or ENO scheme which should be used for temperature. Since such techniques are well known but beyond the scope of this study, we preferred to use the simple clipping approach $T > T_1$ when necessary. From the fine grid computations discussed earlier in this section, the proposed algorithms are accurate enough to represent the near wall complexity of turbulent flows with strong heat transfer. From the coarse grid simulations, these algorithms can handle larger temperature ratios than the methods used in most previous studies.

6. CONCLUSION

Conservation of kinetic energy was specified as an analytical requirement for a proper set of discrete equations in the zero-Mach number limit. The proposed algorithms are fourth-order accurate in space and dissipation-free. A key ingredient is a variable coefficient Poisson equation to solve for pressure. This elliptic operator ensures that the proper constraint is applied to the velocity field at each (sub-)step of the time integration procedure. Noticeably, the divergence-free constraint is recovered in the inviscid limit as required from the low-Mach number approximation to the Navier–Stokes equations. It is shown that this feature is necessary to avoid violating conservation of kinetic energy. A proper discretization of the non-linear terms is also proposed that can handle any density variation and still remains conservative in kinetic energy. The only violation in the kinetic energy conservation appears in the projection step and is related to the non-linear state equation. A fully conservative scheme is obtained if instead an approximate state equation is used.

ACKNOWLEDGMENTS

The author thanks Dr. Oleg Vasilyev for many helpful discussions during the course of this study. C. Pierce and C. McNeil are acknowledged for their remarks on a draft of this manuscript.

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